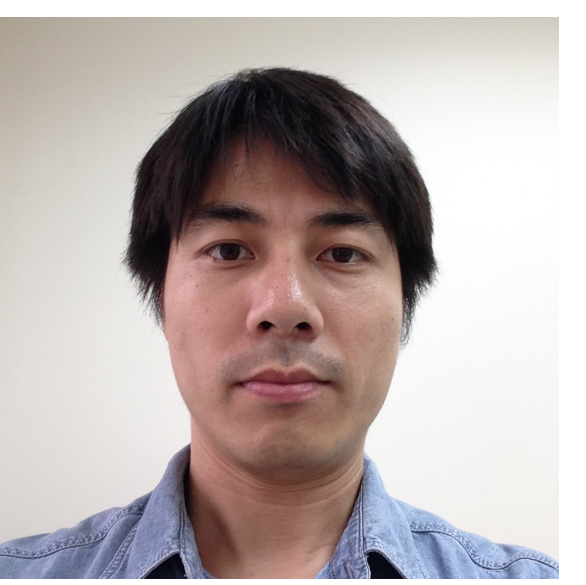


Surface density: a new parameter in the fundamental metallicity relation of star-forming galaxies



Abstract

Tetsuya Hashimoto and Tomotsugu Goto (National Tsing Hua University)

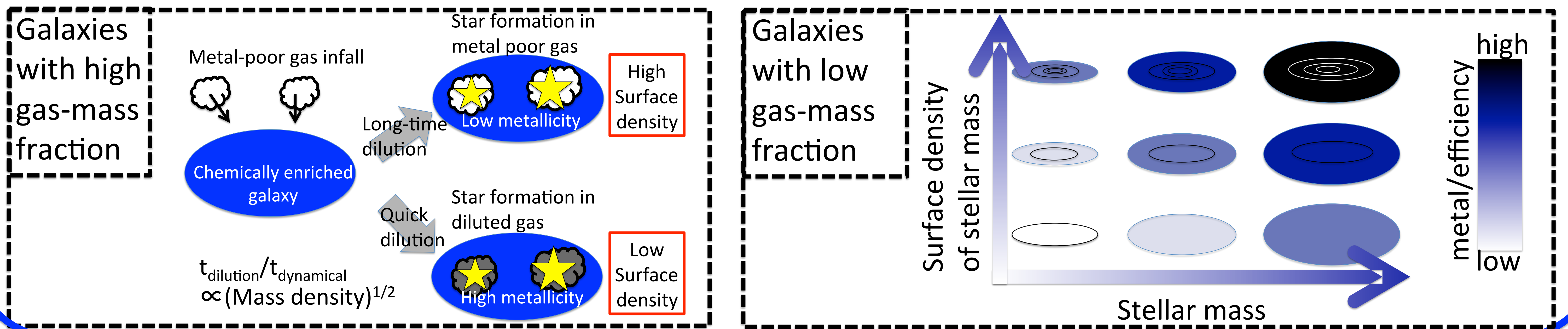
Email: tetsuya@phys.nthu.edu.tw

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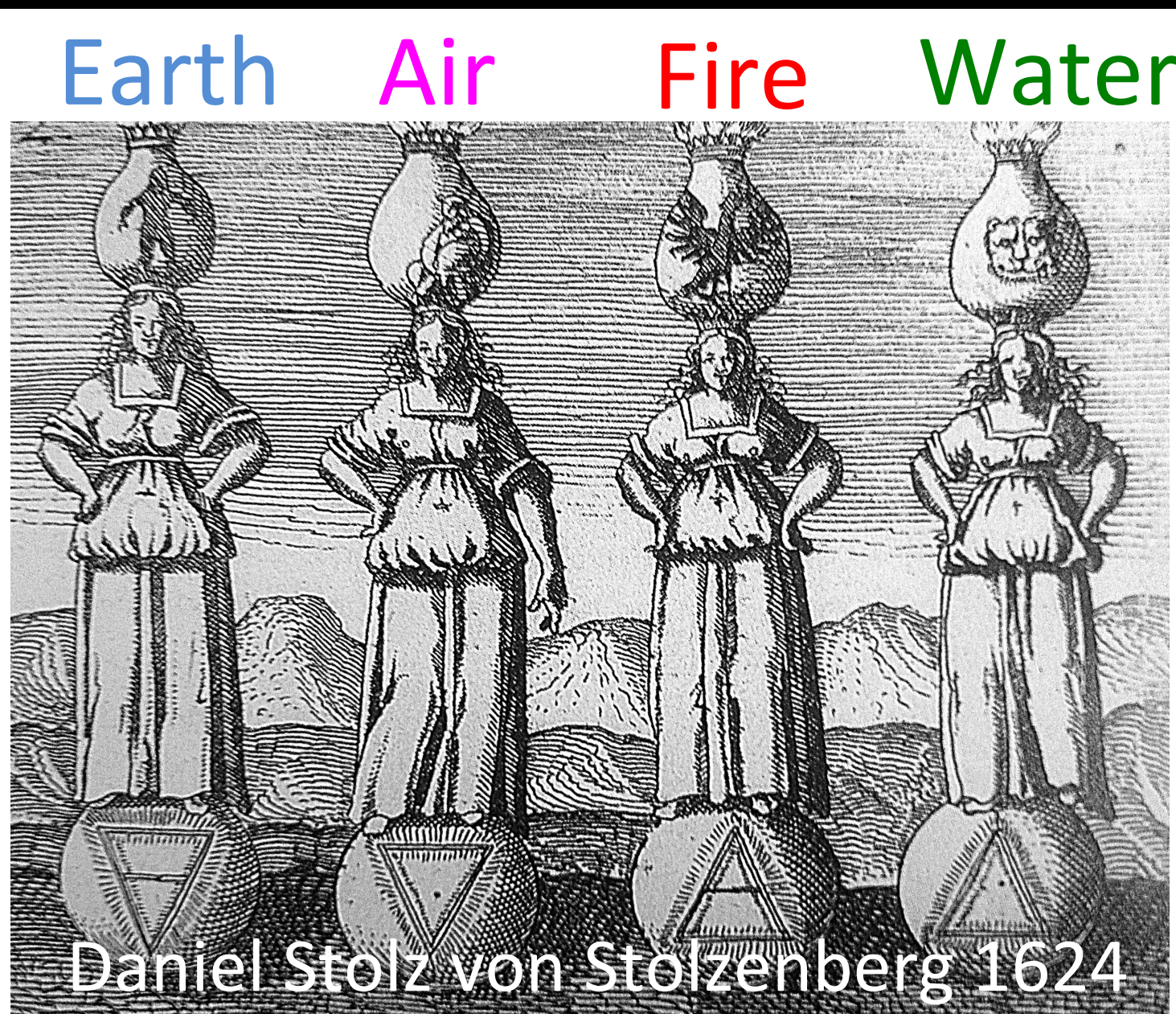
We show here that a fourth parameter, the surface density of stellar mass, reduces the dispersion around the fundamental metallicity relation (FMR). In a principal component analysis of 29 physical parameters of 41,338 star-forming galaxies, the surface density of stellar mass is found to be the fourth most important parameter. The new four-dimensional fundamental relation forms a tighter hypersurface that reduces the metallicity dispersion to 50% of that of the FMR.

Conclusion

Principal component analysis provides us a quantitative way to select most important physical parameters to govern the physics of star-forming galaxies. The negative and positive correlations between metallicity and surface density of stellar mass cause the dispersion around the molecular-gas FMR. The dilution time scale of gas inflow and the star-formation efficiency can explain the observational dependence on surface density.



1. What is the 4th parameter?



Stellar mass SFR
Metallicity



2. Method

Principal component analysis was conducted for 29 physical parameters of SDSS 41,338 SF galaxies. The resulting 'factor loading' is a fractional contribution of each parameter to the total distribution of data points, indicating the quantitative importance. The parameters are listed below in order of the value of factor loading, which are categorized into 6 groups. Highlighted parameter is top one in each group, analyzed further in the right-hand section.

| MASS | METAL | ACTIVITY |
|--|--|--------------------------------------|
| M_* ①Stellar mass | Z_{metal} ④Metallicity | M_{H_2} ③Molecular gas mass |
| M_i i-band abs. mag | M_{metal} Metal mass | SFR Star-formation rate |
| M_z z-band abs. mag | A_v Dust extinction | M_u u-band abs. mag |
| M_r r-band abs. mag | | sSFR Specific SFR |
| M_g g-band abs. mag | | M_{HI} HI gas mass |
| M_{virial} Virial mass | | g-r Color |
| σ Velocity dispersion | | D4000 4000Å break |
| M_{igas} Ionized gas mass | | EW _{Hα} Hα equivalent width |
| SIZE/MORPHOLOGY | ENVIRONMENT | OTHER |
| ΣM_* ②Surface density of M_* | M_{halo} ⑥Dark matter halo mass | q ④Ionization parameter |
| r_{half} Half light radius | δ_5 Galaxy number density | z Redshift |
| ΣSFR Surface density of SFR | | n_e Electron density |
| r_{disk} Disk radius | | |
| B/T Bulge-total fraction | | |

Labeled number is ranking of the factor loading among top-one parameters in 6 groups.

3. Results

The combination of stellar mass, metallicity, molecular-gas mass, and surface density of stellar mass shows the tightest fundamental relation, indicating that the surface density of stellar mass is the 4th most important parameter (top panel). The reduction of the dispersion around the relation is predominant in metallicity axis (middle panel). The correlations between metallicity and the 4th parameter, which depends on the gas-mass fraction, are attributed to the dispersion around the 3D relation (bottom panel).

